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A CUPID $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometer tested in the CROSS underground facility

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ABSTRACT: A scintillating bolometer based on a large cubic $\text{Li}_2^{100}\text{MoO}_4$ crystal (45 mm side) and a Ge wafer (scintillation detector) has been operated in the CROSS cryogenic facility at the Canfranc underground laboratory in Spain. The dual-readout detector is a prototype of the technology that will be used in the next-generation $0\nu 2\beta$ experiment CUPID. The measurements were performed at 18 and 12 mK temperature in a pulse tube dilution refrigerator. This setup utilizes the same technology as the CUORE cryostat that will host CUPID and so represents an accurate estimation of the expected performance. The $\text{Li}_2^{100}\text{MoO}_4$ bolometer shows a high energy resolution of 6 keV FWHM at the 2615 keV γ line. The detection of scintillation light for each event triggered by the $\text{Li}_2^{100}\text{MoO}_4$ bolometer allowed for a full separation ($\sim 8\sigma$) between $\gamma(\beta)$ and α events above 2 MeV. The $\text{Li}_2^{100}\text{MoO}_4$ crystal also shows a high internal radiopurity with ^{228}Th and ^{226}Ra activities of less than 3 and 8 $\mu\text{Bq/kg}$, respectively. Taking also into account the advantage of a more compact and massive detector array, which can be made of cubic-shaped crystals (compared to the cylindrical ones), this test demonstrates the great potential of cubic $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers for high-sensitivity searches for the ^{100}Mo $0\nu 2\beta$ decay in CROSS and CUPID projects.

KEYWORDS: Cryogenic detectors; Double-beta decay detectors; Particle identification methods; Scintillators, scintillation and light emission processes (solid, gas and liquid scintillators)

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1 Introduction

Neutrinoless double-beta ($0\nu 2\beta$) decay is a unique probe of new physics beyond the Standard Model [1, 2] and the observation of this process, suggested about 80 years ago but not yet detected (in contrast to two-neutrino double-beta ($2\nu 2\beta$) decay [3]), would conclusively demonstrate lepton number violation and the Majorana nature of neutrinos (i.e. a particle that is equal to its own anti-particle).

The bolometric technology, which relies on the use of low-temperature calorimeters acting simultaneously as a 2β source and a detector, is among the few experimental approaches providing world-leading sensitivity to $0\nu 2\beta$ decay to-date [2]. In addition to high detection efficiency of the “ 2β source = detector” technique, bolometers offer high energy resolution, scalability to a large detector mass via arrays of modules, and the possibility to use different and radiopure materials containing the most promising 2β isotopes (e.g. see [4–6]). Additionally, recent technological advances have demonstrated the ability to do particle identification [4–6], allowing for a reduction of backgrounds in the signal region of interest by multiple orders of magnitude.

Bolometric techniques for $0\nu 2\beta$ decay searches have been developed for about 30 years and have resulted in the first tonne-scale bolometric experiment CUORE (Cryogenic Underground Observatory of Rare Events) [7]. CUORE has been in operation at the Gran Sasso underground laboratory (Italy) since 2017, searching for $0\nu 2\beta$ decay of ¹³⁰Te (Q -value of the 2β transition, $Q_{2\beta}$, is 2528 keV [8]). In spite of this extraordinary achievement, the CUORE $0\nu 2\beta$ sensitivity is limited by a background ($\sim 10^{-2}$ counts/yr/kg/keV) coming from alpha decays at surfaces despite the highly radiopure materials used for the detector construction. This is due to the use of pure thermal detectors based on tellurium dioxide crystals (TeO₂; 34% of ¹³⁰Te in natural tellurium [9]) which have the same bolometric response irrespective of the type of the particle interaction [10].

CUPID (CUORE Upgrade with Particle IDentification) is a proposed next-generation $0\nu 2\beta$ bolometric experiment [11], which will reuse the CUORE infrastructure for the operation of a similar-scale isotopically enriched detector with a background $\sim 10^{-4}$ counts/yr/kg/keV in the region of interest, thus probing $0\nu 2\beta$ decay in so-called “zero-background” conditions. The suppression of the alpha-induced background to a negligible level (i.e. 99.9% of alpha events rejection), while keeping almost 100% of the signal efficiency, is required for particle identification technology. The detector performance is expected to be similar to CUORE and predecessors, with a 5 keV FWHM at $Q_{2\beta}$ as a goal. The activities of ^{226}Ra , ^{228}Th , and ^{232}Th in the enriched bolometers are required to be less than 10 $\mu\text{Bq/kg}$, making the contribution of the U/Th crystal bulk activity to be below $\sim 10^{-4}$ counts/yr/kg/keV. The total bulk radioactivity of the crystals should not exceed the mBq/kg level to avoid impacting the detector operation and background with pile-ups [5, 12–14].

Four isotopes, ^{82}Se ($Q_{2\beta} = 2998$ keV [8]), ^{100}Mo (3034 keV [8]), ^{116}Cd (2813 keV [8]) and ^{130}Te , were considered in the CUPID R&D program [15] as isotopes of interest to be embedded in the CUPID detector for the following reasons:

- The $0\nu 2\beta$ decay energy of these isotopes (except for ^{130}Te) is greater than 2.6 MeV, the end-point of the most energetic intense natural γ -ray radiation;
- Enrichment is available at a large amount and reasonable cost;
- Compounds containing these isotopes can be grown into single crystals usable for cryogenic applications;
- Some of Se-, Mo-, or Cd-containing crystals are also reasonably efficient low-temperature scintillators. The detection of scintillation light using an auxiliary optical bolometer in coincidences with the measurement of particle-induced energy release in the scintillating absorber is a viable tool for particle identification. This technique can also be applied for poorly or non-scintillating crystals, as TeO_2 , to detect Cherenkov radiation allowing particle identification (however, more performing light detectors are demanded to detect a tiny signal).

Efficient alpha background rejection has been demonstrated with detectors containing each of these isotopes [15–19]. This paves the way for a future study of $0\nu 2\beta$ across multiple isotopes [20] in case a discovery is made. Based on performance and cost, CUPID selected ^{100}Mo embedded in lithium molybdate (Li_2MoO_4) scintillating crystals [21].

The technology of ^{100}Mo -enriched lithium molybdate ($\text{Li}_2^{100}\text{MoO}_4$) scintillating bolometers has been recently developed within the LUMINEU project and it provides [18, 22]:

- A know-how for the mass production of high-quality large radiopure crystals with only few % losses of the enriched material;
- The fabrication of a detector module (which can be easily mount into array) with energy resolution comparable to that of TeO_2 bolometers, but with a significantly higher α rejection efficiency (e.g. see in [5]).

Excellent performances of $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers based on cylindrical crystals ($\varnothing 44 \times 45$ mm, ~ 0.21 kg, $\sim 97\%$ enrichment in ^{100}Mo) have been demonstrated in single-module

and 4-crystal-array tests of LUMINEU [5, 18, 23] at the Gran Sasso and Modane (LSM; France) underground laboratories. These results have been recently confirmed by the CUPID-Mo experiment [24–26] on the scale of a 20-detector array operated at the LSM. Furthermore, the crystal production protocol, adopted by LUMINEU and CUPID-Mo, has been used for the fabrication of 32 $\text{Li}_2^{100}\text{MoO}_4$ crystals 0.28 kg each (the average enrichment in ^{100}Mo is 97.7(3)%) for the CROSS (Cryogenic Rare-event Observatory with Surface Sensitivity) $0\nu 2\beta$ experiment [27].

CROSS, considered as a part of CUPID R&D, is a project aiming at the development of $\text{Li}_2^{100}\text{MoO}_4$ and $^{130}\text{TeO}_2$ surface-coated bolometers capable of identifying a near surface particle interaction via pulse-shape analysis [27]. A key ingredient of the CROSS technology is crystal-surface coating with a superconducting material to modify the signal pulse-shape for an event occurring at its proximity. The CROSS detector performance and radiopurity should be in compliance with CUPID requirements. The feasibility of a highly-efficient identification of near-surface α interactions has recently been demonstrated in multiple tests of the CROSS prototypes [27–30]. A final validation of the technology is planned to be realized as a $0\nu 2\beta$ experiment with at least 32 $\text{Li}_2^{100}\text{MoO}_4$ bolometers (the addition of the 20 crystals from CUPID-Mo are now also in consideration), hosted in a dedicated cryostat at the Canfranc underground laboratory (Spain). The sensitivity of this medium-scale demonstrator [27] is expected to be on the level of the leading $0\nu 2\beta$ experiments, which have masses larger by a factor 10–100.

In contrast to LUMINEU and CUPID-Mo, CROSS is going to use cubic $\text{Li}_2^{100}\text{MoO}_4$ elements with a 45 mm side. The choice of a cubic shape is driven by the possibility to realize a more compact array structure, which allows to deploy a $\sim 30\%$ higher isotope mass in the available experimental volume, and yields an enhanced efficiency in rejecting background-like events that release energy in neighboring crystals (coincidences). Indeed, a volume (i.e. mass) of a cylindrical crystal with a diameter and height equal to the side of the cubic one is almost 30% less (similar to CUPID-Mo vs. CROSS crystals). It is also evident that the efficiency of coincidences between larger, particularly neighbor, detectors would be increased too. The CROSS development of an array of cubic $\text{Li}_2^{100}\text{MoO}_4$ bolometers is also an important benchmark for the design of the final CUPID structure, initially considered to be based on cylindrical ($\varnothing 50 \times 50$ mm) $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers [21]. Before using the crystals in the CROSS and CUPID $0\nu 2\beta$ experiments, it is necessary to perform low-temperature test(s) to demonstrate that:

- Bolometric and spectrometric performances of cubic-shaped $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers, operated in modern pulse-tube cryostats with possible vibration disturbances, are similar to those of cylindrical $\text{Li}_2^{100}\text{MoO}_4$ detectors tested in dry and/or wet dilution refrigerators;
- Scintillation light yield of the cubic-shaped and cylindrical crystals is similar, thus providing a highly efficient particle identification;
- Radioactive contamination of cubic $\text{Li}_2^{100}\text{MoO}_4$ crystals is compatible to that of cylindrical ones.

With these goals in mind, we realized a first investigation of a scintillating bolometer based on a large-volume ($\sim 90 \text{ cm}^3$) cubic-shaped $\text{Li}_2^{100}\text{MoO}_4$ crystal, described in the present paper. This study is undertaken as part of both the CROSS and CUPID R&D programs.

2 Detector construction and operation

2.1 $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometer fabrication

We construct a dual-readout cryogenic particle detector from a primary scintillating absorber and a light detector. The $\text{Li}_2^{100}\text{MoO}_4$ absorber consists of a $45 \times 45 \times 45$ mm crystal ($\sim 98\%$ enrichment in ^{100}Mo) of mass 279.42 g. We randomly chose the sample from the batch of 32 identical crystals. The crystals were grown starting from purified ^{100}Mo powder and using the low-temperature-gradient Czochralski technique at the Nikolaev Institute of Inorganic Chemistry (Novosibirsk, Russia). The $\text{Li}_2^{100}\text{MoO}_4$ samples are not perfectly cubic-shaped¹ due to not optimal growing conditions (in particular, the platinum crucible size was not large enough), adapted for the growth of up to $\varnothing 50$ mm crystal boules [18, 22, 24]. A Neutron Transmutation Doped (NTD) Ge thermistor [31] with a size of $3 \times 3 \times 1$ mm and a P-doped Si chip [32] were epoxy-glued on the crystal top. The dependency of the NTD Ge resistance on temperature can be approximated as $R(T) = R_0 \cdot e^{(T_0/T)^{0.5}}$ with the parameters $T_0 \sim 3.8$ K and $R_0 \sim 1.5 \Omega$. The Si chip is used as a resistive element to periodically inject constant energy pulses used for off-line stabilization of the bolometric response [33]. The crystal holder is made from copper to host cubic crystals of up to 5 cm side and optical bolometers at their top and/or bottom [34–36]. As for the light detector (LD) we use a SiO-coated Ge wafer of 44 mm diameter and 0.175 mm thickness, instrumented with a $3 \times 1 \times 1$ mm NTD. The LD was mounted in the copper holder near the $\text{Li}_2^{100}\text{MoO}_4$ detector (LMO).

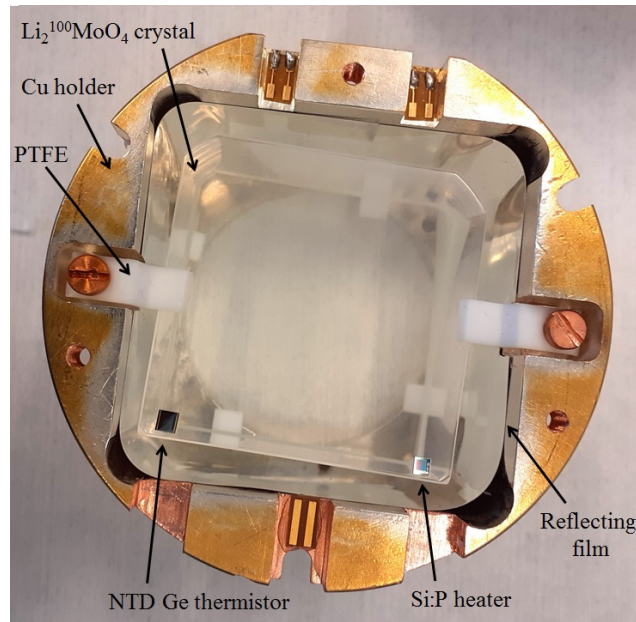


Figure 1. A photograph of the partially assembled $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometer; the construction elements (see text) are labeled. A $\varnothing 45$ mm hole at the bottom of the holder, visible in transparency, acts as an entrance window for the $\text{Li}_2^{100}\text{MoO}_4$ scintillation light to be registered by a bolometric Ge light detector.

The crystal is fixed inside the copper holder by PTFE (polytetrafluoroethylene) pieces, as seen in figure 1. The PTFE supports act also as thermal contacts to the heat sink of the cryostat. The Cu

¹There are edge chamfers on the crystals, see figure 1.

holder of the detector is coated with Au to avoid oxidation, while the internal part of the holder is also coated with Ag to improve light reflection. The $\text{Li}_2^{100}\text{MoO}_4$ crystal inside the Cu housing is surrounded with a Vikuiti™ reflecting film, which is the same used in LUMINEU and CUPID-Mo. The NTD Ge is wire-bonded with Au wires, while the heater is bonded with Al wires. A $^{238}\text{U}/^{234}\text{U}$ source is placed on the holder's top cap. The source was obtained by depositing an uranium-containing liquid drop on a thin copper substrate that was then dried by evaporation. Part of the alpha particles (as well as nuclear recoils) emitted are degraded in energy. The LD is fabricated in the same way using a dedicated Cu holder, three PTFE elements, and an NTD Ge sensor glued. A ^{55}Fe X-ray source is placed close to the LD to irradiate the Ge surface opposite to the $\text{Li}_2^{100}\text{MoO}_4$ absorber.

2.2 Low-temperature underground measurements

We tested the detector in the CROSS Cryogenic Underground (C2U) facility [37], in operation at the Canfranc laboratory (Spain) since April 2019. The cryostat is placed inside a Faraday cage with acoustic isolation, formerly used by the ROSEBUD dark matter experiment [38]. The set-up operates a pulse tube (Cryomech PT415) based dilution refrigerator, developed by CryoConcept (France), which is also assisted by Ultra Quiet Technology™ (UQT) to mitigate vibrations [39]. During the cryostat commissioning, it was found the UQT to efficiently reduce vibrations in the vertical direction, but not as much horizontally [37] resulting in a noise excess affecting the bolometric performance [19]. Thus, the hybrid bolometer was spring-suspended from the detector plate. In order to reduce the environmental background, the cryostat is surrounded externally by a 25 cm thick low-radioactivity lead shield. Moreover, the detector volume inside the cryostat is shielded from the dilution unit and cryostat upper parts with a 13 cm thick disk made of sandwiched lead and copper (120 kg total mass). The shielding of the set-up has not been completed yet, in particular an anti-radon Plexiglas box (to be flushed with a deradonized air) and a muon veto will be installed soon.

The signal readout is based on a low-noise room-temperature DC front-end electronics [40] tracing back to the Cuoricino experiment. The data acquisition (DAQ) is a new design candidate for CUPID [41] and consists of two 12-channel boards with a programmable 6-pole Bessel-Thomson anti-aliasing filter and integrated 24-bit ADC. A cut-off frequency of the low-pass filter can be set from 24 Hz up to 2.5 kHz. With the 24 bit ADC resolution, the input noise is not limited by the ADC even with the lowest gain value set at a programmable-gain amplifier (PGA). An additional advantage of such ADC resolution is that the PGA stage can be made much simpler or removed, with less power consumption, cost, and space. The sampling rate up to 25 kS/s can be set (250 kS/s with half of channels). The ADC-digitized continuous data are readout by an external FPGA (field-programmable gate array) module and then transferred to a personal computer via Ethernet. The DAQ control is done with the help of a MATLAB-based graphical user interface program. The monitoring on-line of the data quality is realized as a LabVIEW application.

We collected data from the end of December 2019 until the beginning of April 2020. The measurements were performed at temperatures 18 and 12 mK. We periodically calibrated the LMO by inserting a thoriated tungsten wire inside the lead shield. We chose the working points for both operational temperatures to be a few nA current on the NTD sensor resulting in a few M Ω resistance. The data are sampled continuously at a 2 kS/s sampling rate, and the full data stream is written to disk for offline analysis. The Bessel-Thomson cut-off frequency was set at 300 Hz, as a compromise between the bandwidth of the LMO (slow) and LD (relatively fast).

Table 1. Performance of a scintillating bolometer based on the $\varnothing 44$ mm Ge light detector coupled to the 45 mm side $\text{Li}_2^{100}\text{MoO}_4$ cubic-shaped scintillator. We report the detectors rise and decay times, the signal amplitude per unit of deposited energy, the energy resolution (FWHM) of the baseline after the optimum filter, at 5.9 keV X-ray of ^{55}Mn (LD), and at 2615 keV γ quanta of ^{208}Tl (LMO). We skip the computation of FWHM at 2615 keV for the 18 mK dataset due to poor statistics of the γ peak. Particle identification parameters (defined in section 3.2) as light yield for $\gamma(\beta)$ s $LY_{\gamma(\beta)}$ and a quenching factor for α particles QF_{α} , as well as the discrimination power between α and $\gamma(\beta)$ distributions $DP_{\alpha/\gamma(\beta)}$ for events selected in the 2.0–5.1 MeV energy range are also quoted.

Channel	Parameter	18 mK	12 mK
LD	Rise time (ms)	1.7	2.8
	Decay time (ms)	9.2	8.6
	Signal ($\mu\text{V}/\text{keV}$)	1.20	1.44
	FWHM (keV) at baseline	0.300(1)	0.210(1)
	FWHM (keV) at 5.9 keV X-ray	0.282(5)	0.315(4)
LMO	Rise time (ms)	18	25
	Decay time (ms)	150	160
	Signal ($\mu\text{V}/\text{keV}$)	0.017	0.036
	FWHM (keV) at baseline	4.2(2)	2.5(1)
	FWHM (keV) at 2615 keV γ	–	6.0(5)
LMO	$LY_{\gamma(\beta)}$ (keV/MeV)	0.635(2)	0.638(1)
& LD	QF_{α} (^{210}Po)	0.192(1)	0.199(4)
	$DP_{\alpha/\gamma(\beta)}$	7.4(4)	7.9(1)

We used around-3-week-long stable periods of data for the analysis at each regulated temperature, not affected by external events (e.g. power cuts). We select 314 h of physics data at each temperature and 65 and 220 h of the ^{232}Th calibration data at 18 and 12 mK, respectively.

The acquired data are triggered offline to tag discrete energy depositions. The triggered pulses are then processed by the optimum filter technique [42] to evaluate the signal amplitude (i.e. energy) and several pulse-shape parameters. In the reconstruction of the coincidences between the LMO and LD, we account for the LD faster response and correct for its constant time shift with respect to the LMO signal, similarly to the method described in [43].

3 Results

3.1 Detector performance

The performance parameters achieved by the LMO and LD in the 18 and 12 mK tests are listed in table 1. The rise and decay time constants, defined respectively as time intervals of the (10–90)% rising edge and (90–30)% trailing edge relative to the signal maximum, of the LMO are ~ 0.02 and ~ 0.15 s, respectively. We expect the LD response to be faster by an order of magnitude because

of the smaller heat capacity of both the Ge absorber and the NTD thermistor. The time constants obtained are in agreement with the results of previous investigations of similar size LMOs and LDs [18, 23]. It is worth noting that the time response of the bolometric detectors depend on the operation temperature and the sensor polarization (see, e.g., [44]). However, optimization of the detector time response² was out of the scope of the present study.

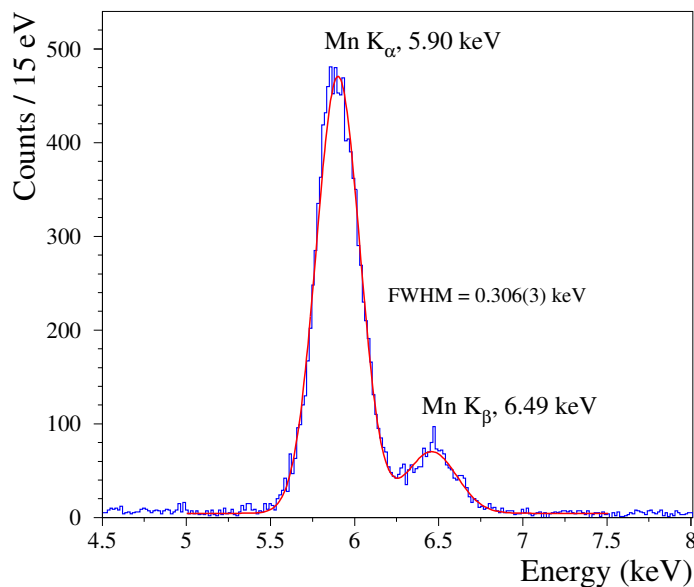


Figure 2. The energy spectrum of the ^{55}Fe X-ray source measured by a 1.4 g Ge bolometric light detector over 913 h (12 and 18 mK data) in the CROSS pulse tube based cryogenic facility at the Canfranc underground laboratory (Spain). A fit to the data by a model assuming a double-Gaussian function and a flat background component is shown by solid (red) line. The energy resolution (FWHM) is quoted for the 5.9 keV X-ray peak of Mn K $_{\alpha}$.

The LD signal amplitude per unit of deposited energy is 1.2 $\mu\text{V}/\text{keV}$ at 18 mK and 1.4 $\mu\text{V}/\text{keV}$ at 12 mK. The LMO signal amplitude is of course inferior, 17 nV/keV at 18 mK and doubles at 12 mK. Since the working points were not optimized to get the highest sensitivity, these results are good but not extraordinary among similar devices [18, 23].

The LD is calibrated with the 5.9 and 6.5 keV X-rays emitted by the ^{55}Fe source. The energy spectrum of the ^{55}Fe source gathered over 913 h of physics and thorium calibration runs is shown in figure 2. The almost fully resolved Mn K $_{\alpha}$ /K $_{\beta}$ doublet is visible thanks to the high LD energy resolution: ≈ 0.3 keV FWHM at 5.9 keV. The baseline noise is 0.2–0.3 keV FWHM, demonstrating a reasonably low threshold. It is worth noting that such devices do not always show a high energy resolution even if characterized by ten(s) eV RMS noise,³ due to the position-dependent response of thin bolometers.

²In particular, to get the fastest response in view of the rejection of random coincidence events induced background in the ^{100}Mo $0\nu 2\beta$ region of interest [12–14].

³For example, the Mn doublet resolution of 0.3–0.5 keV FWHM was measured with LDs made of 30–45 μm thick Ge wafers [18], while the 0.08 keV FWHM resolution was achieved by a 33 g Ge bolometer ($\varnothing 20 \times 20$ mm) characterized by a similar noise level [45].

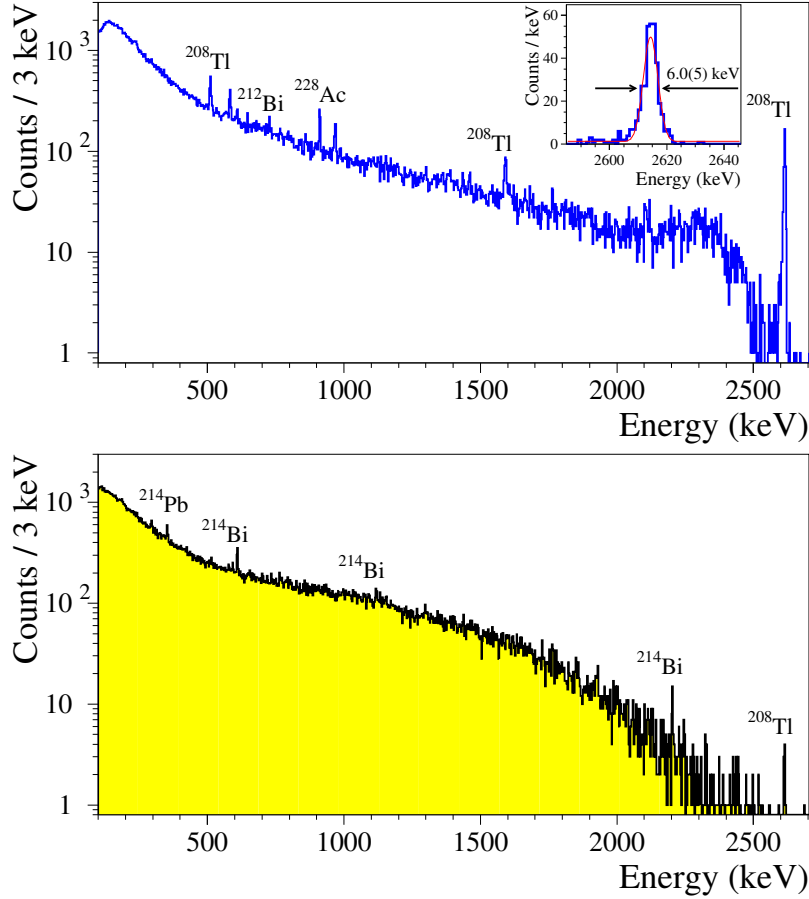


Figure 3. The energy spectra of $\gamma(\beta)$ events accumulated by the LMO over the calibration (285 h; top panel) and physics data (628 h; bottom panel) measurements in the C2U facility at the Canfranc underground laboratory. The most prominent γ peaks are labeled. The contribution of α events to the physics data has been removed with the scintillation light based particle identification (see text). The inset shows the γ peak with energy of 2615 keV in the calibration data together with a fit and the calculated energy resolution (FWHM).

We measured the LMO energy scale and resolution with the most intense gamma peaks in the thorium spectrum, illustrated in figure 3 (top panel). Because of the incomplete shielding, the physics data (figure 3, bottom panel) also exhibit several γ peaks from residual environmental radioactivity (daughters of ^{226}Ra and ^{228}Th sub-chains). The $^{238}\text{U}/^{234}\text{U}$ alpha source also emits β particles from ^{234m}Pa decays ($Q_\beta = 2.27$ MeV [46]), which, together with the ^{100}Mo $2\nu 2\beta$ decays [23], are responsible for the most part of the continuum background above 0.5 MeV, seen in figure 3 (bottom panel).

The energy dependence of the LMO energy resolution is presented in figure 4. The results extracted from physics data are limited by the poor statistics. The detector demonstrates a good energy resolution in a wide energy interval exhibiting a peak width slightly increasing with energy, in agreement with early findings [18, 25]. In particular, we achieved a 6 keV energy resolution (FWHM) for γ -ray quanta of ^{208}Tl with energy 2615 keV, and a 2.5 keV FWHM baseline noise. The resolution at the $Q_{2\beta}$ of ^{100}Mo is expected to be very similar (figure 4). These results are

in agreement with prior measurements for cylindrical LMOs [5, 18, 24], confirming an excellent bolometric performance independent of the crystal shape. It is also evident that a low baseline noise is crucial in obtaining a high energy resolution with a $\text{Li}_2^{100}\text{MoO}_4$ bolometer. It is worth noting, the lowest noise achieved with the LMO is a factor 2–4 worse than the best reported values for large-volume lithium molybdate bolometers [18, 24]. Thus, taking into account sub-optimal noise level of the present study, there is still room for improvement.

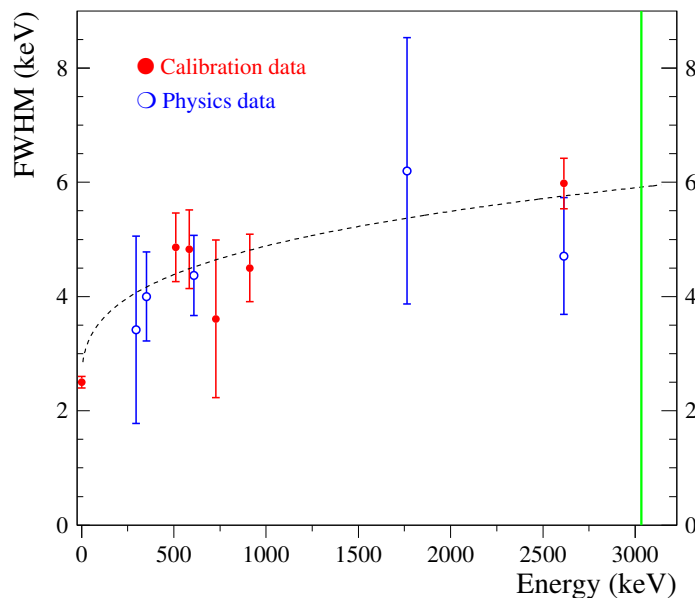


Figure 4. Energy dependence of the 279 g $\text{Li}_2^{100}\text{MoO}_4$ bolometer energy resolution (FWHM) measured in the calibration (filled circles; 12 mK) and physics (open circles; 18 and 12 mK) runs. The fitting curve is shown by the dashed line, while the solid line indicates the $^{100}\text{Mo } Q_{2\beta}$.

3.2 Particle identification capability

Coincidences between the LMO and LD have been used to probe the scintillation based particle identification (PID). For each event triggered by the LMO we calculated a PID parameter, the so-called light yield (LY), defined as the ratio of the LD to LMO measured energy. The LY dependence on particle energy is shown in figure 5, where the population of $\gamma(\beta)$ events is clearly separated from α 's. Such a powerful separation is achieved thanks to the quenching of the scintillation light for α particles with respect to $\gamma(\beta)$'s of the same energy and low noise of the LD. Different ionization properties lead also to a different amplitude measured by the LMO. Figure 5 shows a $\sim 7\%$ increase of an α event energy with respect to the gamma energy scale, in agreement with previous studies of LMO bolometers [18, 24]. This difference, called thermal quenching, hints at a possibility of PID by pulse-shape analysis of the heat channel itself [18], but it is by far less reproducible due to a strong dependence on the noise conditions.⁴

⁴Since the present detector does not have the CROSS technology of the surface coating for PID purpose, we skip an analysis of pulse-shape discrimination of α events.

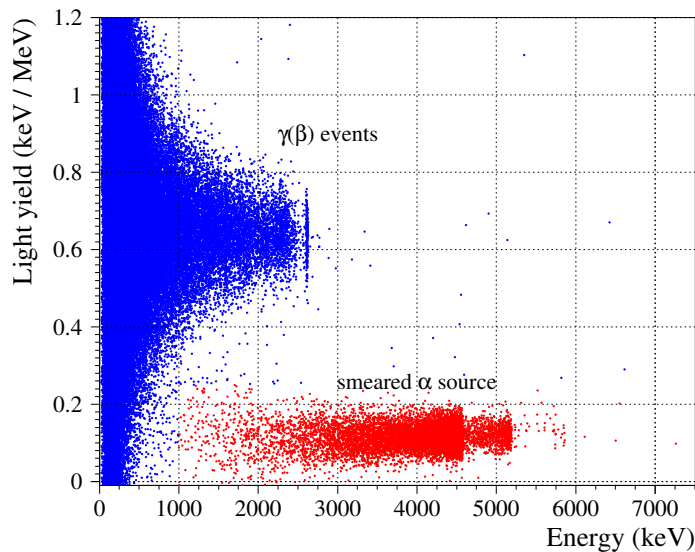


Figure 5. Light yield as a function of energy deposited in the LMO and measured in a 125 h long ^{232}Th calibration at 12 mK. The energy scale is calibrated with γ quanta. The population of $\gamma(\beta)$ events is clearly separated from the α events, mainly originated by the $^{238}\text{U}/^{234}\text{U}$ smeared α source. The α events shown in red were selected above 1 MeV with a LY cut below 0.25 keV/MeV.

In order to investigate a LY -based $\alpha/\gamma(\beta)$ separation close to the ^{100}Mo $0\nu 2\beta$ region of interest (ROI), we selected events within the 2.0–5.1 MeV energy range and a LY interval of 0.4–1.0 keV/MeV for $\gamma(\beta)$'s and 0–0.25 keV/MeV for α 's; both are shown in figure 6. A Gaussian fit to each distribution provides a LY mean value (μ) and standard deviation (σ), which we used to calculate the so-called discrimination power defined as

$$DP_{\alpha/\gamma(\beta)} = |\mu_{\gamma(\beta)} - \mu_{\alpha}| / \sqrt{\sigma_{\gamma(\beta)}^2 + \sigma_{\alpha}^2}. \quad (3.1)$$

The $DP_{\alpha/\gamma(\beta)}$ based on the present data is around 8 (table 1), meaning about $8\sigma_{\alpha}$ of alpha event rejection while keeping almost 100% of $\gamma(\beta)$'s. This rejection power fully satisfies the CUPID goal of identifying 99.9% of alpha particles (corresponding to $DP_{\alpha/\gamma(\beta)} \sim 3.1$).

The LMO light yield for $\gamma(\beta)$ events ($LY_{\gamma(\beta)}$) was found to be 0.64 keV/MeV, similar to one measured with cylindrical shaped LMOs of $\varnothing 44 \times 45$ mm size [18, 24, 26]. However, in the present study the light collection efficiency was affected by the smaller area of the LD (15 cm^2) with respect to the LMO surface facing it (20 cm^2) and by the entrance window of the Cu holder ($\varnothing 45$ mm). Thus, considering only the direct light, the $LY_{\gamma(\beta)}$ is expected to be ~ 0.85 keV/MeV, once a 45 mm side square LD is coupled to the LMO. In case of the use of two LDs, the LY should be roughly doubled, as demonstrated by CUPID-Mo [24, 26]. The quenching factor (QF_{α}) for α particles of ^{210}Po observed in the data (see the next section) is 0.2,⁵ in agreement with the previous data [18, 24, 26].

⁵Such a parameter is typically quoted without the correction from the thermal quenching and we follow that convention here.

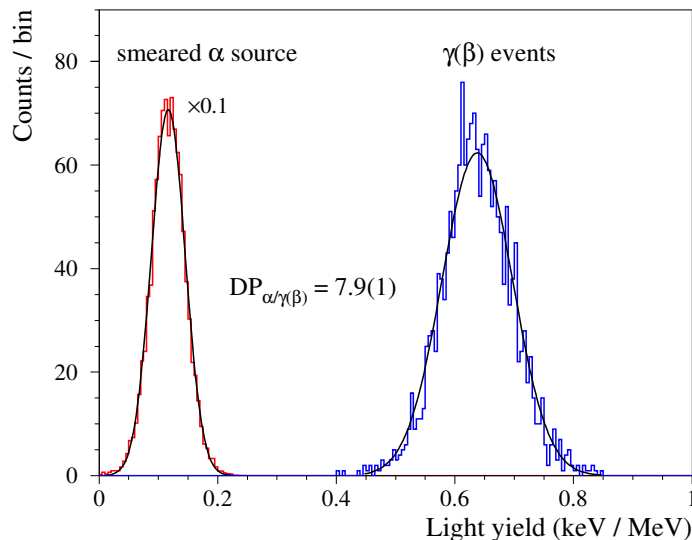


Figure 6. The light yield distributions for α and $\gamma(\beta)$ particles, selected in the 2.0–5.1 MeV energy interval. The former distribution is rescaled by a factor 0.1 to improve the visibility of the later one. A Gaussian fit to each distribution is shown by a solid line. The discrimination power between the two populations is ~ 8 .

3.3 $\text{Li}_2^{100}\text{MoO}_4$ crystal radiopurity

A highly-efficient PID together with a good energy resolution of the LMO operated over four weeks of background measurements allow us to quantify the $\text{Li}_2^{100}\text{MoO}_4$ radiopurity with a high sensitivity. The spectrum of alpha events selected from the physics data (and recalibrated to α energy) is presented in figure 7; the energy interval covers most Q_α -values of radionuclides from the U/Th chains. As it is seen in figure 7, the use of the $^{238}\text{U}/^{234}\text{U}$ smeared α source prevents estimation of the alpha activity of ^{232}Th , ^{238}U and some of the daughters with $Q_\alpha \leq 4.8$ MeV. However, we can investigate a possible contamination by ^{226}Ra and ^{228}Th , which are the most harmful contaminants for $0\nu 2\beta$ searches.

The energy region above 4.8 MeV contains only two peak-like structures both ascribed to ^{210}Po α events, and originated by the ^{210}Pb contamination [47]. A clear peak at 5.4 MeV is induced by the ^{210}Po decays in the $\text{Li}_2^{100}\text{MoO}_4$ crystal bulk with the activity of 80(12) $\mu\text{Bq/kg}$. A ^{210}Po (^{210}Pb) bulk contamination on the level of ten(s)–hundred(s) $\mu\text{Bq/kg}$ is typical for $\text{Li}_2^{100}\text{MoO}_4$ crystals [5, 18, 24, 26]. A broad distribution peaked at 5.3 MeV is caused by the ^{210}Po decay on the surface of materials facing the LMO (a 0.1 MeV energy is taken away by the ^{206}Pb nuclear recoil). The decays of ^{210}Po at surfaces of the detector materials can populate the ^{100}Mo $0\nu 2\beta$ ROI as energy-degraded alpha events, but they can be easily rejected thanks to the efficient PID of $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers. Also, the total rate of ^{210}Po events (~ 0.06 mHz) is rather low to be a notable source of pile-ups, which are of certain concern for slow response thermal detectors [16].

The spectrum shows no other structures, so we can only set limits on other radionuclide contaminations. In order to be conservative, we considered all events within 25 keV of each expected peak location. The background (1 count per 50 keV) was estimated in the 5.65–5.75 and 5.85–5.95 MeV energy intervals, containing no Q_α -value of U/Th radionuclides. The data exhibit

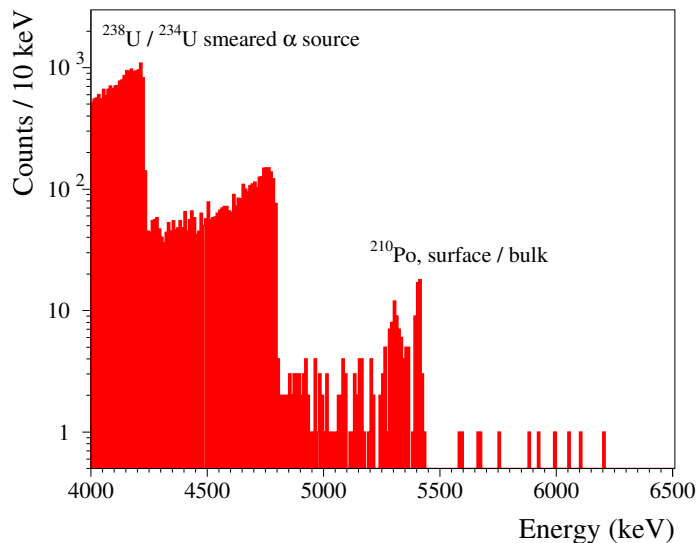


Figure 7. A part of the α energy spectrum accumulated by the LMO operated over 628 h in the CROSS underground cryostat in Canfranc. The energy interval covers the Q_α -values of the most α -active radionuclides from U/Th decay chains. In addition to a dominant contribution from the used $^{238}\text{U}/^{234}\text{U}$ source, the spectrum exhibits only two populations of ^{210}Po originated by external (surface of nearby materials) or internal (crystal bulk) contaminations of the detector.

no events of ^{228}Th ($Q_\alpha = 5520$ keV [8]) and two counts of ^{222}Rn ($Q_\alpha = 5590$ keV [8]; a daughter of ^{226}Ra), thus we place 90 confidence level (C.L.) upper limits of 1.6 and 4.9 counts, respectively [48]. A selection efficiency of 95.7% was estimated using alpha events of the $^{238}\text{U}/^{234}\text{U}$ smeared α source, distributed in the 3.5–4.8 MeV energy interval. Therefore, the activity of ^{228}Th and ^{226}Ra in the $\text{Li}_2^{100}\text{MoO}_4$ crystal bulk is below 3 and 8 $\mu\text{Bq}/\text{kg}$ at 90% C.L., respectively. Taking into account the reasonably short half-lives of ^{228}Ra (5.75 yr [46]) and ^{228}Th (1.91 yr [46]), the limit on the ^{228}Th activity can also represent the ^{232}Th contamination in the crystals, as e.g. seen in [18, 25, 26].

The limits on the ^{228}Th and ^{226}Ra activity in the studied $\text{Li}_2^{100}\text{MoO}_4$ crystal are on the same level as reported by LUMINEU [5, 18], which were obtained by the analysis of a comparable exposure. A significantly larger exposure of the CUPID-Mo experiment [24–26] shows that the level of remaining contaminants in $\text{Li}_2^{100}\text{MoO}_4$ scintillators can be even an order of magnitude lower. As the protocol of the CROSS crystal production was the one adopted by LUMINEU and CUPID-Mo, it seems natural to assume the radiopurity of the CROSS crystals to be similar to that of CUPID-Mo ones. It is worth noting that the ^{228}Th (^{232}Th) and ^{226}Ra contamination on the level of 10 $\mu\text{Bq}/\text{kg}$ is compatible with a background contribution below 10^{-4} counts/yr/kg/keV to the ^{100}Mo $0\nu 2\beta$ ROI, and it is fully acceptable not only for the medium-scale CROSS experiment (with ~ 10 kg detector mass), but also for the tonne-scale extension of CUPID.

4 Conclusions

We report that performance and radiopurity of a scintillating bolometer based on a large cubic-shaped $\text{Li}_2^{100}\text{MoO}_4$ crystal — randomly taken from 32 identical crystals (with a 45 mm side and a mass of 0.28 kg each) of the CROSS $0\nu 2\beta$ project — are similar to those of cylindrical 0.2 kg

$\text{Li}_2^{100}\text{MoO}_4$ bolometers, used in LUMINEU and CUPID-Mo 2β experiments. In particular, the $\text{Li}_2^{100}\text{MoO}_4$ detector energy resolution at 2615 keV γ quanta (as well as its approximation to 3034 keV, the ^{100}Mo 2β decay energy) is 6 keV FWHM. A scintillation light yield of the cubic-shaped crystal, 0.64 keV/MeV for $\gamma(\beta)$'s, is compatible to that of cylindrical crystals. However, the measured light yield is affected by about 30% lower detection surface of the optical bolometer with respect to the nearby $\text{Li}_2^{100}\text{MoO}_4$ crystal face. In spite of sub-optimal light collection, a full separation ($\sim 8\sigma$) between α and $\gamma(\beta)$ events above 2 MeV has been achieved. A high radiopurity of the cubic-shaped $\text{Li}_2^{100}\text{MoO}_4$ crystal was also demonstrated by the present study, where only ^{210}Po is detected with the activity of 80(12) $\mu\text{Bq/kg}$, while the content of ^{228}Th and ^{226}Ra (the most harmful radionuclides from U/Th families for $0\nu 2\beta$ searches) is estimated to be less than 3 and 8 $\mu\text{Bq/kg}$, respectively.

The performed investigation additionally proves the excellent prospects of $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers for high-sensitivity $0\nu 2\beta$ decay searches. The cubic shape of large-volume ($\sim 90\text{ cm}^3$) $\text{Li}_2^{100}\text{MoO}_4$ crystals allows a more compact detector array structure and thus the deployment of a larger isotope mass in the experimental volume, as well as an increased efficiency of multi-site event detection. In view of these results, large-mass ($\sim 0.3\text{ kg}$) radiopure cubic-shaped $\text{Li}_2^{100}\text{MoO}_4$ crystals operated as bolometers satisfy the demands of the CROSS and CUPID projects.

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References

- [1] J.D. Vergados, H. Ejiri and F. Šimkovic, *Neutrinoless double beta decay and neutrino mass*, *Int. J. Mod. Phys. E* **25** (2016) 1630007 [[arXiv:1612.02924](#)].
- [2] M.J. Dolinski, A.W.P. Poon and W. Rodejohann, *Neutrinoless Double-Beta Decay: Status and Prospects*, *Ann. Rev. Nucl. Part. Sci.* **69** (2019) 219 [[arXiv:1902.04097](#)].
- [3] A. Barabash, *Precise Half-Life Values for Two-Neutrino Double- β Decay: 2020 Review*, *Universe* **6** (2020) 159 [[arXiv:2009.14451](#)].
- [4] S. Pirro and P. Mauskopf, *Advances in Bolometer Technology for Fundamental Physics*, *Ann. Rev. Nucl. Part. Sci.* **67** (2017) 161.
- [5] LUMINEU, CUPID-0/Mo and EDELWEISS collaborations, *^{100}Mo -enriched Li_2MoO_4 scintillating bolometers for $0\nu 2\beta$ decay search: From LUMINEU to CUPID-0/Mo projects*, *AIP Conf. Proc.* **1894** (2017) 020017 [[arXiv:1709.07846](#)].
- [6] F. Bellini, *Potentialities of the future technical improvements in the search of rare nuclear decays by bolometers*, *Int. J. Mod. Phys. A* **33** (2018) 1843003 [[arXiv:1802.04260](#)].
- [7] CUORE collaboration, *Improved Limit on Neutrinoless Double-Beta Decay in ^{130}Te with CUORE*, *Phys. Rev. Lett.* **124** (2020) 122501 [[arXiv:1912.10966](#)].
- [8] M. Wang et al., *The AME2016 atomic mass evaluation (II). Tables, graphs and references*, *Chin. Phys. C* **41** (2017) 030003.
- [9] J. Meija et al., *Isotopic compositions of the elements 2013 (IUPAC Technical Report)*, *Pure Appl. Chem.* **88** (2016) 293.
- [10] F. Bellini et al., *Response of a TeO_2 bolometer to α particles*, *2010 JINST* **5** P12005 [[arXiv:1010.2618](#)].
- [11] CUPID collaboration, *CUPID: CUORE (Cryogenic Underground Observatory for Rare Events) Upgrade with Particle Identification*, [arXiv:1504.03599](#).
- [12] D.M. Chernyak, F.A. Danevich, A. Giuliani, E. Olivieri, M. Tenconi and V.I. Tretyak, *Random coincidence of $2\nu 2\beta$ decay events as a background source in bolometric $0\nu 2\beta$ decay experiments*, *Eur. Phys. J. C* **72** (2012) 1989 [[arXiv:1301.4248](#)].
- [13] D.M. Chernyak et al., *Rejection of randomly coinciding events in ZnMoO_4 scintillating bolometers*, *Eur. Phys. J. C* **74** (2014) 2913 [[arXiv:1404.1231](#)].
- [14] D.M. Chernyak et al., *Rejection of randomly coinciding events in $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers using light detectors based on the Neganov-Luke effect*, *Eur. Phys. J. C* **77** (2017) 3 [[arXiv:1606.02287](#)].
- [15] CUPID collaboration, *R & D towards CUPID (CUORE Upgrade with Particle Identification)*, [arXiv:1504.03612](#).
- [16] D. Poda and A. Giuliani, *Low background techniques in bolometers for double-beta decay search*, *Int. J. Mod. Phys. A* **32** (2017) 1743012 [[arXiv:1711.01075](#)].
- [17] CUPID-0 collaboration, *Final result of CUPID-0 phase-I in the search for the ^{82}Se Neutrinoless Double- β Decay*, *Phys. Rev. Lett.* **123** (2019) 032501 [[arXiv:1906.05001](#)].
- [18] E. Armengaud et al., *Development of ^{100}Mo -containing scintillating bolometers for a high-sensitivity neutrinoless double-beta decay search*, *Eur. Phys. J. C* **77** (2017) 785 [[arXiv:1704.01758](#)].

- [19] D.L. Helis et al., *Neutrinoless Double-Beta Decay Searches with Enriched $^{116}\text{CdWO}_4$ Scintillating Bolometers*, *J. Low Temp. Phys.* **199** (2020) 467.
- [20] A. Giuliani, F.A. Danevich and V.I. Tretyak, *A multi-isotope $0\nu 2\beta$ bolometric experiment*, *Eur. Phys. J. C* **78** (2018) 272 [[arXiv:1712.08534](#)].
- [21] CUPID collaboration, *CUPID pre-CDR*, [arXiv:1907.09376](#).
- [22] V.D. Grigorieva et al., *Li_2MoO_4 crystals grown by low thermal gradient Czochralski technique*, *J. Mat. Sci. Eng. B* **7** (2017) 63.
- [23] E. Armengaud et al., *Precise measurement of $2\nu\beta\beta$ decay of ^{100}Mo with the CUPID-Mo detection technology*, *Eur. Phys. J. C* **80** (2020) 674 [[arXiv:1912.07272](#)].
- [24] E. Armengaud et al., *The CUPID-Mo experiment for neutrinoless double-beta decay: performance and prospects*, *Eur. Phys. J. C* **80** (2020) 44 [[arXiv:1909.02994](#)].
- [25] CUPID collaboration, *First data from the CUPID-Mo neutrinoless double beta decay experiment*, *J. Phys. Conf. Ser.* **1468** (2020) 012129 [[arXiv:1911.10426](#)].
- [26] CUPID-Mo collaboration, *Performance of the CUPID-Mo double-beta decay bolometric experiment*, in *XXIX International (online) Conference on Neutrino Physics and Astrophysics (Neutrino 2020)*, June 22–July 02, 2020 [<https://indico.fnal.gov/event/19348/contributions/186385/>].
- [27] CROSS collaboration, *The $0\nu 2\beta$ -decay CROSS experiment: preliminary results and prospects*, *JHEP* **01** (2020) 018 [[arXiv:1906.10233](#)].
- [28] H. Khalife et al., *The CROSS Experiment: Rejecting Surface Events by PSD Induced by Superconducting Films*, *J. Low Temp. Phys.* **199** (2020) 19 [[arXiv:1907.11411](#)].
- [29] CROSS collaboration, *The CROSS experiment: rejecting surface events with PSD*, in *XXIX International (online) Conference on Neutrino Physics and Astrophysics (Neutrino 2020)*, June 22–July 02, 2020 [<https://indico.fnal.gov/event/19348/contributions/186492/>].
- [30] CROSS collaboration, *First results of CROSS underground measurements with massive bolometers*, in *XXIX International (online) Conference on Neutrino Physics and Astrophysics (Neutrino 2020)*, June 22–July 02, 2020 [<https://indico.fnal.gov/event/19348/contributions/186216/>].
- [31] E.E. Haller, *Advanced far-infrared detectors*, *Infrared Phys. Techn.* **35** (1994) 127.
- [32] E. Andreotti et al., *Production, characterization, and selection of the heating elements for the response stabilization of the CUORE bolometers*, *Nucl. Instrum. Meth. A* **664** (2012) 161.
- [33] A. Alessandrello et al., *Methods for response stabilization in bolometers for rare decays*, *Nucl. Instrum. Meth. A* **412** (1998) 454.
- [34] V. Novati, *Sensitivity enhancement of the CUORE experiment via the development of Cherenkov hybrid TeO_2 bolometers*, Ph.D. Thesis, University Paris-Saclay, Orsay, France, November (2018).
- [35] L. Bergé et al., *Complete event-by-event $\alpha/\gamma(\beta)$ separation in a full-size TeO_2 CUORE bolometer by Neganov-Luke-magnified light detection*, *Phys. Rev. C* **97** (2018) 032501 [[arXiv:1710.03459](#)].
- [36] CROSS collaboration, *The CROSS experiment: search for $0\nu 2\beta$ decay with surface sensitive bolometers*, *J. Phys. Conf. Ser.* **1468** (2020) 012147.
- [37] CROSS collaboration, *The new CROSS Cryogenic Underground (C2U) facility: an overview*, in *XXIX International (online) Conference on Neutrino Physics and Astrophysics (Neutrino 2020)*, June 22–July 02, 2020 [<https://indico.fnal.gov/event/19348/contributions/186315/>].

- [38] S. Cebrian et al., *Bolometric WIMP search at Canfranc with different absorbers*, *Astropart. Phys.* **21** (2004) 23.
- [39] E. Olivieri, J. Billard, M. De Jesus, A. Juillard and A. Leder, *Vibrations on pulse tube based Dry Dilution Refrigerators for low noise measurements*, *Nucl. Instrum. Meth. A* **858** (2017) 73 [[arXiv:1703.08957](#)].
- [40] C. Arnaboldi et al., *The programmable front-end system for CUORICINO, an array of large-mass bolometers*, *IEEE Trans. Nucl. Sci.* **49** (2002) 2440.
- [41] P. Carniti, C. Gotti and G. Pessina, *High resolution digitization system for the CROSS experiment*, *J. Low Temp. Phys.* **199** (2019) 833 [[arXiv:1908.11242](#)].
- [42] E. Gatti and P. Manfredi, *Processing the signals from solid-state detectors in elementary-particle physics*, *Riv. Nuovo Cim.* **9** (1986) 1.
- [43] G. Piperno, S. Pirro and M. Vignati, *Optimizing the energy threshold of light detectors coupled to luminescent bolometers*, *2011 JINST* **6** P10005 [[arXiv:1107.5679](#)].
- [44] J.W. Beeman et al., *Characterization of bolometric Light Detectors for rare event searches*, *2013 JINST* **8** P07021 [[arXiv:1304.6289](#)].
- [45] EDELWEISS collaboration, *Searching for low-mass dark matter particles with a massive Ge bolometer operated above-ground*, *Phys. Rev. D* **99** (2019) 082003 [[arXiv:1901.03588](#)].
- [46] G. Audi, F.G. Kondev, M. Wang, W.J. Huang and S. Naimi, *The NUBASE2016 evaluation of nuclear properties*, *Chin. Phys. C* **41** (2017) 030001.
- [47] LUMINEU and EDELWEISS collaborations, *Development and underground test of radiopure ZnMoO₄ scintillating bolometers for the LUMINEU 0ν2β project*, *2015 JINST* **10** P05007.
- [48] G.J. Feldman and R.D. Cousins, *Unified approach to the classical statistical analysis of small signals*, *Phys. Rev. D* **57** (1998) 3873 [[physics/9711021](#)].